

Effective Field Theories with Celestial Duals

Anastasia Volovich
Brown University

Mago, Ren, Spradlin, Yelleshpur Srikant



2111.11356 and 2206.08322

Introduction

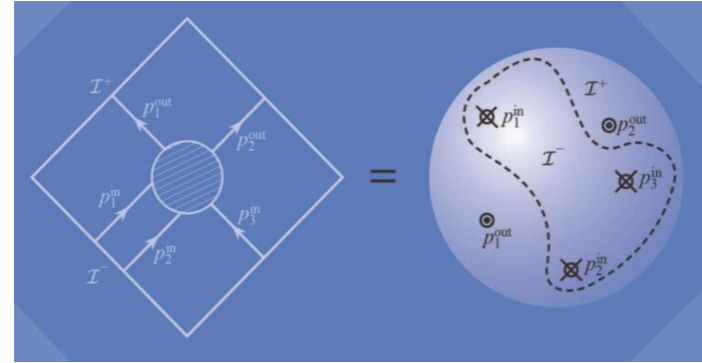
Celestial Holography program is interesting both for understanding holographic description of flat space and for exploring properties of scattering amplitudes.

Stieberger's review talk

I will discuss necessary conditions for a 4d theory to have a celestial dual with an associative OPE.

Celestial Holography

Massless amplitude in 4d



$$\begin{array}{c}
 \mathcal{A}_n(p_i, s_i) \rightarrow \int \prod_i dw_i w_i^{\Delta_i - 1} \mathcal{A}_n(p_i, s_i) \rightarrow \\
 \begin{array}{c}
 \uparrow \text{momentum} \quad \uparrow \text{helicity} \\
 p = w(1 + z\bar{z}, z + \bar{z}, i(z - \bar{z}), 1 - z\bar{z})
 \end{array} \\
 \begin{array}{c}
 \uparrow \text{conformal weights} \\
 s = h - \bar{h}, \Delta = h + \bar{h}
 \end{array}
 \end{array}$$

$$\rightarrow \langle \mathcal{O}_{h_1, \bar{h}_1}(z_1, \bar{z}_1) \dots \mathcal{O}_{h_n, \bar{h}_n}(z_n, \bar{z}_n) \rangle$$

Correlation function in 2d
Celestial Conformal Field Theory

Pasterski, Shao, Strominger

Plan

- Review $w_{1+\infty}$ -algebra in CCFT
Strominger
- Deform $w_{1+\infty}$ due to non-minimal couplings
MRSV
- Jacobi identity \Rightarrow constraints on couplings
MRSV
- Implications for amplitudes in theories
w/ celestial duals
RSSV
- Conclusion / open questions

Graviton OPEs

Two positive helicity gravitons

$$G_{\Delta_1}^+(z, \bar{z}) G_{\Delta_2}^+(0, 0) \sim -\frac{1}{2} \frac{\kappa_{2,2-2}}{z} \sum_{n=0}^{\infty} B(\Delta_1 - 1 + n, \Delta_2 - 1) \frac{\bar{z}^{n+1}}{n!} \partial^n G_{\Delta_1 + \Delta_2}^+$$

$B(a, b) = \frac{\Gamma(a)\Gamma(b)}{\Gamma(a+b)}$

$\sqrt{G_N}$

- from collinear limit of amplitudes
Fan, Fotopoulos, Stieberger, Taylor, Zhu
- from symmetries
Pate, Raclariu, Strominger, Yuan

Conformally soft gravitons

- Noting that OPE coefficients B have poles, define *conformally soft graviton operators*

$$H^K(z, \bar{z}) = \lim_{\Delta \rightarrow K} (\Delta - K) G_{\Delta}^+(z, \bar{z}) \quad K=2, 1, 0, -1, \dots$$

Pate, Raclariu, Strominger

- Mode expansion

$$H^K(z, \bar{z}) = \sum_{n=\frac{K-2}{2}}^{\frac{z-K}{2}} \frac{H_n^K(z)}{\bar{z}^{n+\frac{K-2}{2}}}$$

- We can work out OPE of $H_n^K(z)$

Soft current algebra

- Defining the commutator

$$[A, B](z) = \oint_z \frac{dw}{2\pi i} A(w) B(z)$$

- Guevara, Himwich, Pate, Strominger worked out soft current algebra for gravity \rightarrow long expressions

- Strominger defined $w_n^p = \frac{1}{\mathcal{K}_{2,2,-2}} (p-n-1)!(p+n-1)! H_n^{-2p+4}$ "light-transform" and showed that commutation relations are

$$[w_m^p, w_n^q] = (m(q-1) - n(p-1)) w_{m+n}^{p+q-2} \quad \begin{array}{l} p = 1, \frac{3}{2}, 2, \frac{5}{2}, \dots \\ 1-p \leq m \leq p-1 \end{array}$$

w-algebra

$$[W_m^p, W_n^q] = (m(q-1) - n(p-1)) W_{m+n}^{p+q-2}$$

$$p = 1, \frac{3}{2}, 2, \frac{5}{2}, \dots$$
$$1-p \leq m \leq p-1$$

" $W_{1+\infty}$ "

Penrose 1976

Bakas 1989

Pope Romans Shen 1990

Klebanov Polyakov 1991

$$p=2 \Rightarrow c=0 \quad \text{Virasoro} \quad \checkmark$$

- There is a similar story for gluons

$$[S_n^{q,a}, S_m^{p,b}] = -i f_{abc} S_{n+m}^{q+p-1,c}$$

$$[W_m^p, S_n^{q,a}] = (m(q-1) + n(p-1)) S_{m+n}^{p+q-2,a}$$

Deformed OPE

- Himwich, Pate, Singh considered corrections to OPE from non-minimal couplings.

$$\begin{array}{c} 2^{++} \\ | \\ \text{---} \circ \text{---} 3^{--} \\ / \quad \backslash \\ 1^{++} \end{array} = \kappa_{2,2,-2} \frac{[12]^6}{[23]^2 [31]^2} \quad \mathcal{R}$$

$$\begin{array}{c} 2^{++} \\ | \\ \text{---} \circ \text{---} 3^{++} \\ / \quad \backslash \\ 1^{++} \end{array} = \kappa_{2,2,2} [12]^2 [23]^2 [13]^2 \quad \mathcal{R}^3$$

$\hookrightarrow \kappa_{h_1, h_2, h_3} = 3\text{pt coupling for helicities } h_1, h_2, h_3$

$$G_{\Delta_1}^+(z, \bar{z}) G_{\Delta_2}^+(0, 0) \sim \frac{1}{z} \left(\kappa_{2,2,-2} B(\Delta_1 - 1, \Delta_2 - 1) \bar{z} G_{\Delta_1 + \Delta_2}^+ + \kappa_{2,2,2} B(\Delta_1 + 3, \Delta_2 + 3) \bar{z}^5 G_{\Delta_1 + \Delta_2 + 4}^{+-} \right)$$

Deformed w-algebra

• The algebra becomes

$$[W_{m_1}^{q_1, 2}, W_{m_2}^{q_2, 2}] = - \frac{\kappa_{2,2,2} (m_1(q_1-1) - m_2(q_2-1))}{2} W_{m_1+m_2}^{q_1+q_2-2, 2}$$

$$- \frac{\kappa_{2,2,2}}{2} N(q_1, q_2, m_1, m_2, 5) W_{m_1+m_2}^{q_1+q_2-6, 2}$$

$$N(q_1, q_2, m_1, m_2, p) = \sum_{n=0}^p (-1)^{p-n} \binom{p}{n} [m_1+q_1-1]_{p-n} [-m_1+q_1-1]_n$$

$$[m_2+q_2-1]_n [-m_2+q_2-1]_{p-n}$$

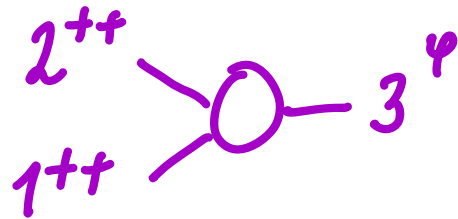
where $[a]_n \equiv a(a-1)\dots(a-n+1)$

Pope, Romans, Shen 1990

Jacobi identity

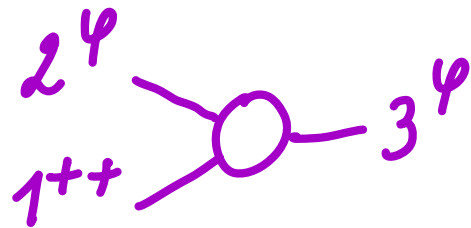
$$[[W_{m_1}^{q,2}, W_{m_2}^{q,2}], W_{m_3}^{q,2}] + \text{cyclic} \neq 0 \quad ?$$

Add scalars



$$\kappa_{0,2,2} [12]^4$$

$$\Psi R^2$$



$$\kappa_{0,0,2} \frac{[12]^2 [13]^2}{[23]^2}$$

$$R \Psi^2$$

Jacobi identity requires

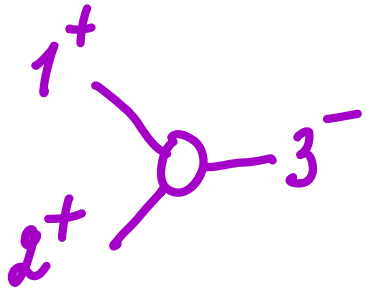
$$(\kappa_{-2,2,2} - \kappa_{0,0,2}) \kappa_{0,2,2} = 0$$

$$(\kappa_{-2,2,2} - \kappa_{0,0,2}) \kappa_{0,0,2} = 0$$

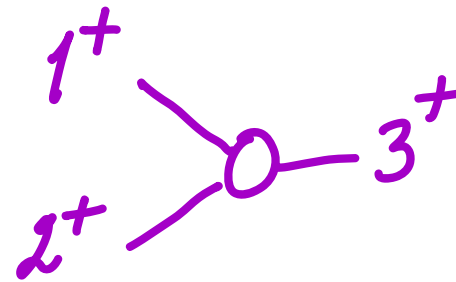
$$3 \kappa_{0,2,2}^2 = 10 \kappa_{-2,2,2} \kappa_{2,2,2}$$

MRSV

Include gluons

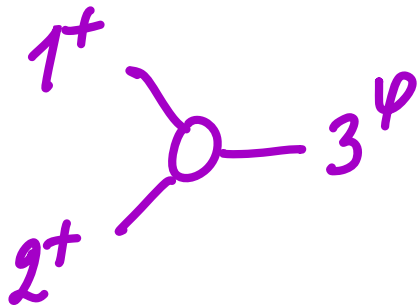


$$\alpha_{-1,1,1}$$



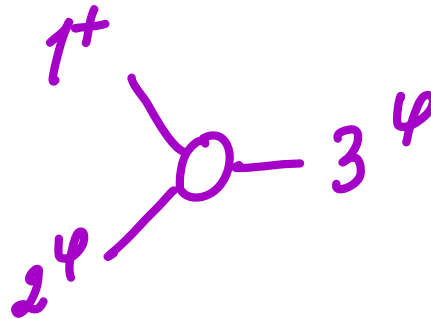
$$\alpha_{1,1,1}$$

$$F^3$$



$$\alpha_{0,1,1}$$

$$\psi F^2$$



$$\alpha_{0,0,1}$$

$$\psi^2 F$$

Constraints

$$\alpha_{0,1,1}^2 = 2 \alpha_{-1,1,1} \alpha_{1,1,1}$$

\swarrow YM \swarrow F3

$$\alpha_{-1,1,2} \alpha_{1,1,1} = 3 \alpha_{1,1,2} \alpha_{-1,1,1}$$

MRSV

Relation to W-algebra

Pope, Shen, Romans constructed $W_{1+\infty}$ by deforming $W_{1+\infty}$

$$[V_{m_1}^{q_1}, V_{m_2}^{q_2}] = \sum_{r=0}^{\infty} \lambda^{2r} g_{2r}^{q_1 q_2}(m_1, m_2; \mu) V_{m_1+m_2}^{q_1+q_2-2r-2} + \lambda^{2(q_1-2)} c_{q_1}(m_1; \mu) \delta^{q_1, q_2} \delta_{m_1+m_2, 0}$$

- $\mu = -3/16$, $g_{2r}^{q_1 q_2} = \frac{1}{2(2r+1)} N(q_1, q_2, m_1, m_2, 2r+1)$

- $c_q = 0$

- wedge subalgebra $q = 1, 3/2, 2, \dots$ $1-q \leq m \leq q-1$

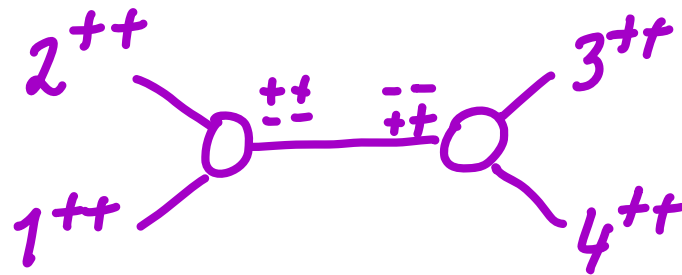
- $[V_{m_1}^{q_1, s_1}, V_{m_2}^{q_2, s_2}] = \sum_{r=0}^{\infty} \frac{\lambda^{2r}}{2(2r+1)!} N(q_1, q_2, m_1, m_2, 2r+1) \cdot V_{m_1+m_2}^{q_1+q_2-2r-2s, s_1+s_2-2r-2}$

OUR ALGEBRA $s = 0, \pm 2$

Implications for Amplitudes 1

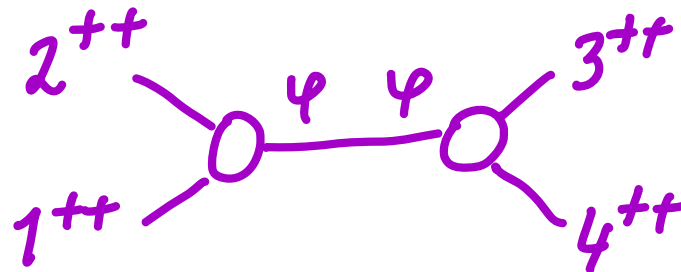
$$A(1^{++}, 2^{++}, 3^{++}, 4^{++}) =$$

Broedel, Dixon



graviton
exchange

+



scalar
exchange

+ permutations

$$= (10 \kappa_{2,2,2} \kappa_{2,2,2} - 3 \kappa_{0,2,2}^2) \#$$

$$= 0$$

VANISHES IF WE
IMPOSE OUR CONSTRAINT!

RSSV

Implications for Amplitudes 2

- Similarly $A(1^{++}, 2^{++}, 3^{++}, 4^{\psi}) = 0$ on support of our constraints
 $A(1^+, 2^+, 3^+, 4^+) = 0$
 $A(1^+, 2^+, 3^+, 4^{\psi}) = 0$
- We showed that all (all-line shift constructible) four-point amplitudes vanish on the support of our constraints
- These amplitudes are constructed from purely holomorphic / antiholomorphic vertices.

Implications for Amplitudes 3

- Higher-point amplitudes

$$A(1^+, 2^+, 3^+, 4^+, 5^+) \neq 0$$

- all-line shift constructible

- has a term involving $A(1^+, 2^+, 3^+)A(1^-, 2^+, 3^+, 4^+)$

↑ includes
holomorphic
+
antiholomorphic
vertices

- All-line shift constructible vanish only if they don't receive contributions from any channels other than those involving only pure holomorphic/antiholomorphic vertices as building blocks

OPE associativity RSSV

- We can derive the coupling constraints directly from the associativity of the OPEs (w/o W -algebra)

$$\left(\underset{2 \rightarrow 3}{\text{Res}} \underset{1 \rightarrow 2}{\text{Res}} - \underset{1 \rightarrow 3}{\text{Res}} \underset{2 \rightarrow 3}{\text{Res}} + \underset{2 \rightarrow 3}{\text{Res}} \underset{1 \rightarrow 3}{\text{Res}} \right) \langle \mathcal{O}_{\Delta_1}(z_1, \bar{z}_1) \cdots \mathcal{O}_{\Delta_n}(z_n, \bar{z}_n) \rangle = 0$$

- The correlator is Mellin transform of the amplitude.
- We can check associativity of celestial OPE by evaluating double residues directly on the amplitude.

$$\sum_{S_I} A_3(\tilde{\lambda}_1^{s_1}, \tilde{\lambda}_2^{s_2}, \tilde{\lambda}_I^{-s_I}) \frac{1}{[12][34]} A_3(\tilde{\lambda}_I^{s_I}, \tilde{\lambda}_3^{s_3}, \tilde{\lambda}_4^{s_4}) + \text{permutations} = 0$$

- This immediately gives coupling constraints equations.

The following are equivalent:

1. **Soft** Jacobi/double residue condition (wedge part)

2. **Hard celestial** double residue condition

3. **Hard momentum space** double residue condition

4. ~~\exists~~ 4-pt amplitudes w/ angle bracket weight -1

e.g.
$$\frac{[12]\langle 12 \rangle}{\langle 13 \rangle^2}$$

→ Purely an amplitudes statement.
Ready to use straight out of box!



Adam Ball arXiv:2211.09151

Self-dual gravity

- In quantum self-dual gravity $W_{1+\infty}$ algebra persists w/o corrections (one-loop exact all-plus amplitudes Bern et al!)
Ball, Narayanan, Salzer, Strominger
- Moyal deformation of self-dual gravity leads to $W_{1+\infty}$ algebra Monteiro
Bu, Heuveline, Skinner
- Celestial $W_{1+\infty}$ symmetries from twistor space Adamo, Mason, Sharma
- Celestial dual of self-dual YM form factors don't have associative chiral algebra unless one adds an 'axion' Castello, Paquette

Summary

- worked out deformations of $w_{1+\infty}$ due to non-minimal couplings
- Jacobi identity / OPE associativity implies constraints on coupling
- worked out implications of these constraints on amplitudes

What theories satisfy our constraints ?

- Chiral higher-spin theory

$$\kappa_{s_1, s_2, s_3} \sim \frac{(l_P)^{s_1 + s_2 + s_3 - 1}}{\Gamma(s_1 + s_2 + s_3)} \quad s_1 + s_2 + s_3 > 0$$

Metsaev '91

- String theory compactified to 4d?